

Building blocks for a statistically advanced daily temperature reconstruction system

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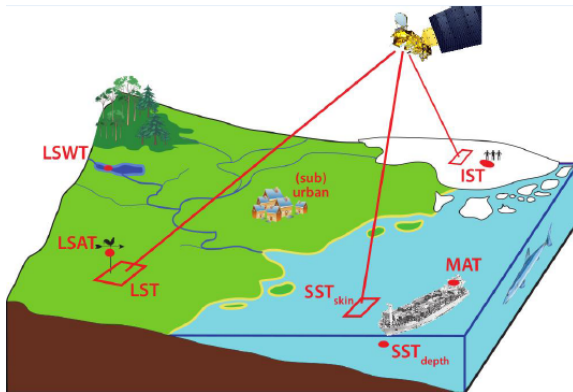


EUSTACE

EU Surface Temperatures for All Corners of Earth

EUSTACE goal:

Daily estimates of surface air temperature since 1850 across the globe by combining surface and satellite data using novel statistical techniques.



From Merchant et al, *Geosci. Instrum. Method. Data Syst.*, 2, 305-321, 2013



Statistical model and method building blocks

Basic system components

- ▶ Multiple *observation sources*, with complex error *uncertainty structure*
- ▶ Temperature *processes on different spatial and temporal scales*
 - ▶ Seasonal
 - ▶ Slow climate processes
 - ▶ Medium-scale variability
 - ▶ Daily
- ▶ *Vast model size* ($\sim 10^{11}$ unknowns); need computationally efficient tools
- ▶ Hierarchical statistical model structure based on Gaussian processes
 - ▶ Stochastic PDEs translates to sparse precisions in *Gaussian Markov random fields*
- ▶ *Propagated uncertainty* via a Bayesian approach
 - ▶ Dependence structure parameters
 - ▶ Spatio-temporal process priors
 - ▶ Observation models
- ▶ Goals:
 - ▶ a *best estimate*,
 - ▶ a *collection of samples*, and
 - ▶ more precise (and accurate) *uncertainty estimates*.

Matérn driven heat equation on the sphere

The iterated heat equation is a simple non-separable space-time SPDE family:

$$(\kappa^2 - \Delta)^{\gamma/2} \left[\phi \frac{\partial}{\partial t} + (\kappa^2 - \Delta)^{\alpha/2} \right]^{\beta} x(\mathbf{s}, t) = \mathcal{W}(\mathbf{s}, t) / \tau$$

For constant parameters, $x(\mathbf{s}, t)$ has spatial Matérn covariance (for each t).

Discrete domain Gaussian Markov random fields (GMRFs)

$\mathbf{x} = (x_1, \dots, x_n) \sim \mathcal{N}(\boldsymbol{\mu}, \mathbf{Q}^{-1})$ is Markov with respect to a neighbourhood structure $\{\mathcal{N}_i, i = 1, \dots, n\}$ if $Q_{ij} = 0$ whenever $j \notin \mathcal{N}_i \cup i$.

- ▶ Project the SPDE solution space onto local basis functions: random Markov dependent basis weights (Lindgren et al, 2011).

A finite element approximation has structure

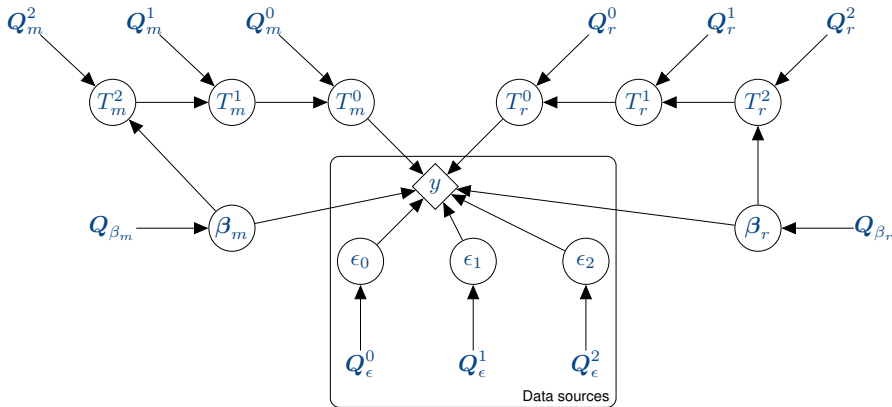
$$x(\mathbf{s}, t) = \sum_{i,j} \psi_i^{[s]}(\mathbf{s}) \psi_j^{[t]}(t) x_{ij}, \quad \mathbf{x} \sim \mathcal{N}(\mathbf{0}, \mathbf{Q}^{-1}), \quad \mathbf{Q} = \sum_{k=0}^{\alpha+\beta+\gamma} \mathbf{M}_k^{[t]} \otimes \mathbf{M}_k^{[s]}$$

even, e.g., if the spatial scale parameter κ is spatially varying.



Partial hierarchical representation

Observations of *mean, max, min*. Model *mean and range*.



Conditional specifications, e.g.

$$(T_m^0 | T_m^1, Q_m^0) \sim \mathcal{N}(T_m^1, Q_m^0^{-1})$$

$$T_r^0 = \exp(T_r^1) G^{-1}[U_r^0(\mathbf{s}, t)], \quad U_r^0 \sim \mathcal{N}(\mathbf{0}, Q_r^0^{-1})$$

Standardised observation uncertainty models

- ▶ Each data source may have complicated dependence structure
- ▶ To facilitate information blending, use a common error term structure

Common satellite derived data error model framework

The observational&calibration errors are modelled as three error components:

- ▶ independent (ϵ_0),
- ▶ spatially and/or temporally correlated (ϵ_1), and
- ▶ systematic (ϵ_2),

with distributions determined by the uncertainty information from satellite calibration models.

E.g., $y_i = T_m(\mathbf{s}_i, t_i) + \epsilon_0(\mathbf{s}_i, t_i) + \epsilon_1(\mathbf{s}_i, t_i) + \epsilon_2(\mathbf{s}_i, t_i)$

In practice, each data source might have several different components of each type; independent components can be merged, but not necessarily correlated or systematic components.

Station observation & homogenisation model

Daily means

For station k at day t_i ,

$$y_m^{k,i} = T_m(\mathbf{s}_k, t_i) + \sum_{j=1}^{J_k} H_j^k(t_i) e_m^{k,j} + \epsilon_m^{k,i},$$

where $H_j^k(t)$ are temporal step functions, $e_m^{k,j}$ are latent bias variables, and $\epsilon_m^{k,i}$ are independent measurement and discretisation errors.

Daily mean/max/min

For station k at day t_i ,

$$y_m^{k,i} = T_m(\mathbf{s}_k, t_i) + \tilde{H}_m^k(t_i) + \epsilon_m^{k,i},$$

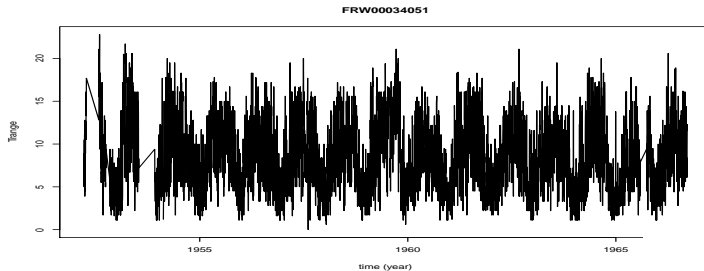
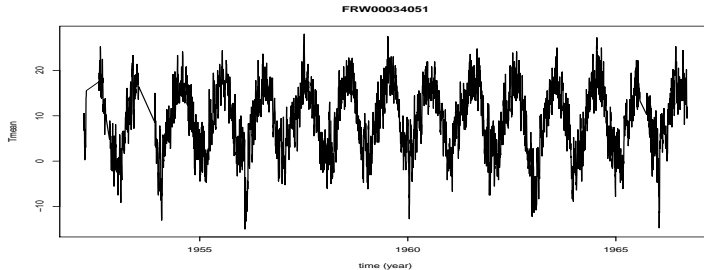
$$y_x^{k,i} = T_m(\mathbf{s}_k, t_i) + \frac{\exp[\tilde{H}_r^k(t_i)]}{2} T_r(\mathbf{s}_k, t_i) + \epsilon_x^{k,i},$$

$$y_n^{k,i} = T_m(\mathbf{s}_k, t_i) - \frac{\exp[\tilde{H}_r^k(t_i)]}{2} T_r(\mathbf{s}_k, t_i) + \epsilon_n^{k,i},$$

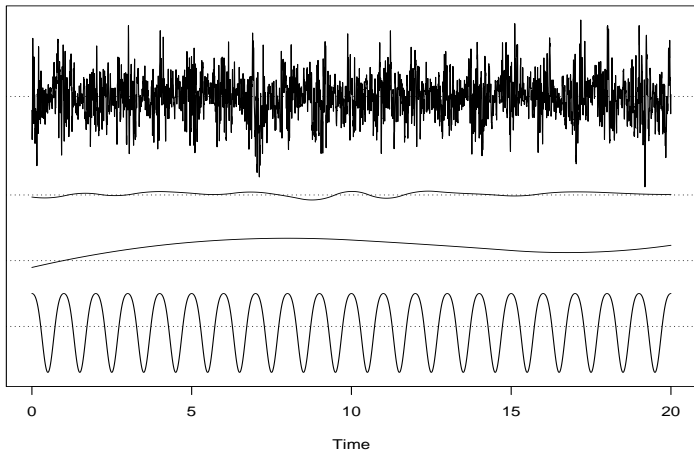
where \tilde{H} : are the total bias correction variables for each observation.

Observed data

Observed daily T_{mean} and T_{range} for station FRW00034051

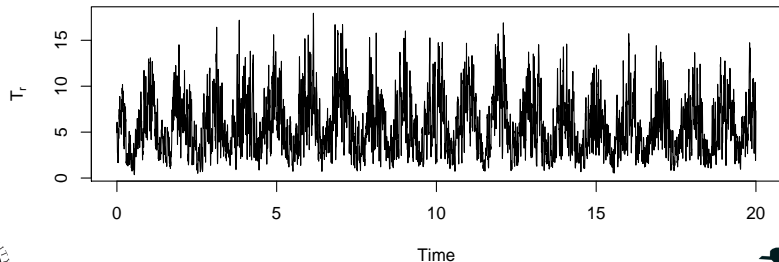
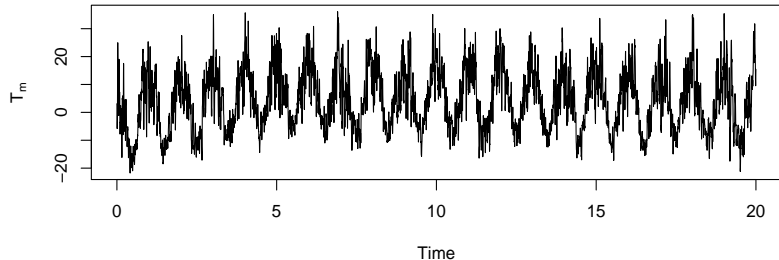


Multiscale model component samples



Combined model samples for T_m and T_r

(Proof of concept; no actual data was involved in this figure)



Modelling non-Gaussian quantities

Power tail quantile (POQ) model

The quantile function $F_{\theta}^{-1}(p)$, $p \in [0, 1]$, is defined through a quantile blend of left- and right-tailed generalised Pareto distributions:

$$f_{\theta}^{-}(p) = \begin{cases} \frac{1-(2p)^{-\theta}}{2\theta}, & \theta \neq 0, \\ \frac{1}{2} \log(2p), & \theta = 0, \end{cases}$$

$$f_{\theta}^{+}(p) = -f_{\theta}^{-}(1-p) = \begin{cases} \frac{(2(1-p))^{-\theta}-1}{2\theta}, & \theta \neq 0, \\ -\frac{1}{2} \log(2(1-p)), & \theta = 0. \end{cases}$$

$$F_{\theta}^{-1}(p) = \theta_0 + \frac{\tau}{2} [(1-\gamma)f_{\theta_3}^{-}(p) + (1+\gamma)f_{\theta_4}^{+}(p)].$$

The parameters $\theta = (\theta_0, \theta_1 = \log \tau, \theta_2 = \text{logit}[(\gamma+1)/2], \theta_3, \theta_4)$ control the median, spread/scale, skewness, and the left and right tail shape.

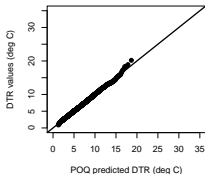
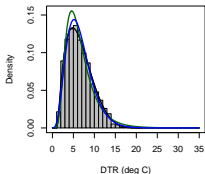
This model is also known as the *five parameter lambda model* (Gilchrist, 2000).

Copula transformation: $G^{-1}[u(s, t)] = F_{\theta(s, t)}^{-1}\{\Phi[u(s, t)]\}$

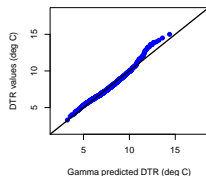
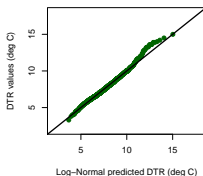
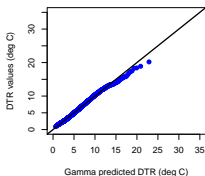
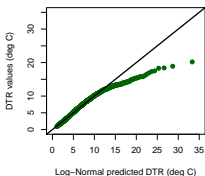
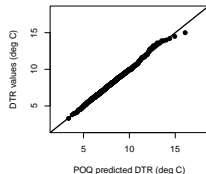
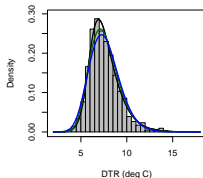


Diurnal range distributions

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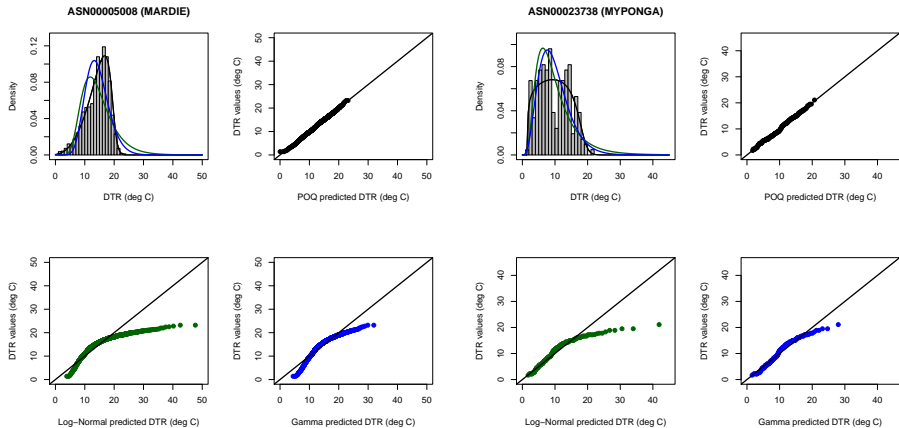


SP000060040 (LANZAROTE/AEROPUERTC)



For these stations, POQ does a slightly better job than a Gamma distribution.

Diurnal range distributions

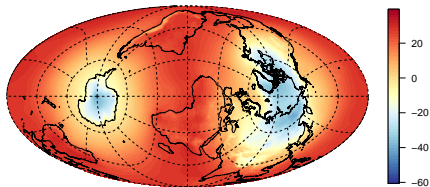


For these stations only POQ comes close to representing the distributions.

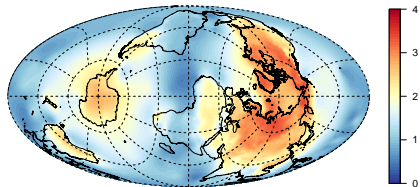
Note: Some of the mixture-like distribution shapes may be an effect of unmodeled station inhomogeneities.

Estimates of median & scale for T_m and T_r

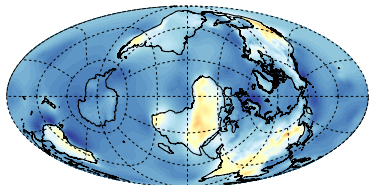
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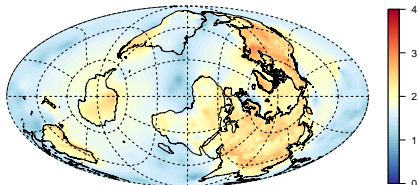
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February climatology

(Preliminary estimates, using only in-situ land station data)

Linearised inference

All Spatio-temporal latent random processes combined into $\mathbf{x} = (\mathbf{u}, \boldsymbol{\beta}, \mathbf{b})$, with joint expectation $\boldsymbol{\mu}_x$ and precision \mathbf{Q}_x :

$$(\mathbf{x} \mid \boldsymbol{\theta}) \sim \mathcal{N}(\boldsymbol{\mu}_x, \mathbf{Q}_x^{-1}) \quad (\text{Prior})$$

$$(\mathbf{y} \mid \mathbf{x}, \boldsymbol{\theta}) \sim \mathcal{N}(h(\mathbf{A}\mathbf{x}), \mathbf{Q}_{y|x}^{-1}) \quad (\text{Observations})$$

$$p(\mathbf{x} \mid \mathbf{y}, \boldsymbol{\theta}) \propto p(\mathbf{x} \mid \boldsymbol{\theta}) p(\mathbf{y} \mid \mathbf{x}, \boldsymbol{\theta}) \quad (\text{Conditional posterior})$$

Non-linear and/or non-Gaussian observations

For a non-linear $h(\mathbf{A}\mathbf{x})$ with Jacobian \mathbf{J} at $\mathbf{x} = \tilde{\boldsymbol{\mu}}$, iterate:

$$(\mathbf{x} \mid \mathbf{y}, \boldsymbol{\theta}) \stackrel{\text{approx}}{\sim} \mathcal{N}(\tilde{\boldsymbol{\mu}}, \tilde{\mathbf{Q}}^{-1}) \quad (\text{Approximate conditional posterior})$$

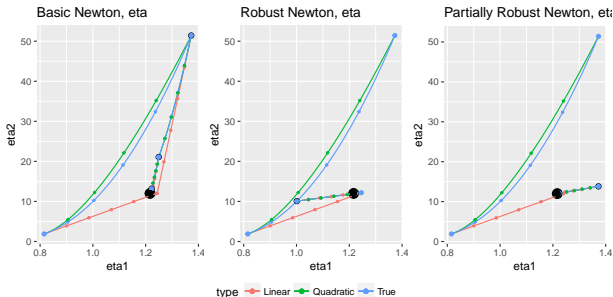
$$\tilde{\mathbf{Q}} = \mathbf{Q}_x + \mathbf{J}^\top \mathbf{Q}_{y|x} \mathbf{J}$$

$$\tilde{\boldsymbol{\mu}}' = \tilde{\boldsymbol{\mu}} + a \tilde{\mathbf{Q}}^{-1} \left\{ \mathbf{J}^\top \mathbf{Q}_{y|x} [\mathbf{y} - h(\mathbf{A}\tilde{\boldsymbol{\mu}})] - \mathbf{Q}_x (\tilde{\boldsymbol{\mu}} - \boldsymbol{\mu}_x) \right\}$$

for some $a > 0$ chosen by line-search.

Iterative solutions for $\sim 10^{11}$ latent variables

- ▶ Nonlinear Newton iteration with robust line-search



- ▶ Preconditioned conjugate gradient (PCG) iteration for $Q(\mu - \hat{\mu}) = r = b - Q\hat{\mu}$
- ▶ Local and multiscale approximations for preconditioning: $M^{-1}Q \approx I$
- ▶ Sampling with PCG: $Q(x - \hat{\mu}) = Lw$
Requires only a rectangular pseudo-Cholesky factorisation $LL^T = Q$.
Possible due to the kronecker product sum precision structure.

Summary

Not covered in this talk:

- ▶ Pure conditional block updates risk getting stuck; need for convergence acceleration
- ▶ Overlapping space-time blocks for preconditioning
- ▶ Non-stationary random field parameter estimation
- ▶ Direct&iterative variance calculations to eliminate or reduce Monte Carlo error in the reconstruction uncertainties
- ▶ Fast approximate handling of correlated error components

Summary:

- ▶ Challenging statistical problem, in both size and complexity
- ▶ Approximate calculation techniques allows some of the complexity to be handled with reasonable computational resources
- ▶ Close collaboration between climate scientistis, statisticians, and software engineers is essential



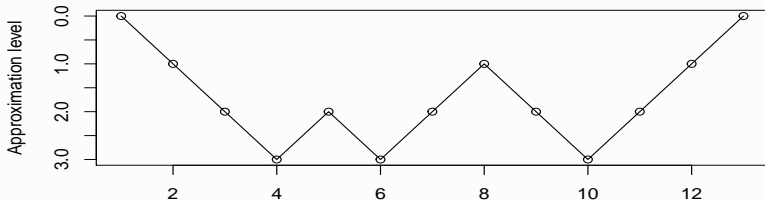
Overlapping blocks and multigrid

Overlapping block preconditioning

Let D_k^\top be a restriction matrix to subdomain Ω_k , and let W_k be a diagonal weight matrix. Then an additive Schwarz preconditioner is

$$M^{-1}x = \sum_{k=1}^K W_k D_k (D_k^\top Q D_k)^{-1} D_k^\top W_k x$$

Multigrid and/or approximate multiscale Schur complements



Complications: Schur complements vs conditional block updating

Variance calculations

Sparse partial inverse: Takahashi recursions postprocesses Cholesky

Takahashi recursions compute \mathbf{S} such that $\mathbf{S}_{ij} = (\mathbf{Q}^{-1})_{ij}$ for all $Q_{ij} \neq 0$.
Postprocessing of the (sparse) Cholesky factor.

Basic Rao-Blackwellisation of sample estimators

Let $\mathbf{x}^{(j)}$ be samples from a Gaussian posterior and let $\mathbf{a}^\top \mathbf{x}$ be a linear combination of interest. Then, for any subdomain $\Omega_k \subset \Omega$,

$$\mathbb{E}(\mathbf{a}^\top \mathbf{x}) = \mathbb{E} [\mathbb{E}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*})] \approx \frac{1}{J} \sum_{j=1}^J \mathbb{E}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*}^{(j)})$$

$$\begin{aligned} \text{Var}(\mathbf{a}^\top \mathbf{x}) &= \mathbb{E} [\text{Var}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*})] + \text{Var} [\mathbb{E}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*})] \\ &\approx \text{Var}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*}^j) + \frac{1}{J} \sum_{j=1}^J [\mathbb{E}(\mathbf{a}^\top \mathbf{x} \mid \mathbf{x}_{\Omega_k^*}^{(j)}) - \mathbb{E}(\mathbf{a}^\top \mathbf{x})]^2 \end{aligned}$$

Efficient if $\mathbf{a}\mathbf{a}^\top$ sparsity matches \mathbf{S} for each subdomain.

Converting Gaussian to POQ

A POQ copula model

A spatio-temporally dependent Gaussian field $u(\mathbf{s}, t)$ with expectation $\mathbf{0}$ and variance $\mathbf{1}$ can be transformed into a POQ field by

$$\tilde{u}(\mathbf{s}, t) = G^{-1}[u(\mathbf{s}, t)] = F_{\theta(\mathbf{s}, t)}^{-1}[\Phi(u(\mathbf{s}, t))],$$

where the parameters can vary with space and time.

Due to the large size of the problem, we estimate parameters in a two-step procedure:

1. Estimate seasonal POQ and temporal covariance parameters for separate time series
2. With a basic spatial-seasonal random field prior, find the posterior mean parameter field